

Energy Conversion by Semiconductor Devices

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U.S. Department of Energy
Energy Efficiency
and Renewable Energy

The History of Solar

Solar technology isn't new. Its history spans from the 7th Century B.C. to today. We started out concentrating the sun's heat with glass and mirrors to light fires. Today, we have everything from solar-powered buildings to solar-powered vehicles.

Here you can learn more about the milestones in the historical development of solar technology, century by century, and year by year. You can also glimpse the future.

This timeline lists the milestones in the historical development of solar technology from the 7th Century B.C. to the 1200s A.D.

7th Century B.C.
Magnifying glass used to concentrate sun's rays to make fire and to burn ants.

3rd Century B.C.
Greeks and Romans use burning mirrors to light torches for religious purposes.

2nd Century B.C.
As early as 212 BC, the Greek scientist, Archimedes, used the reflective properties of bronze shields to focus sunlight and to set fire to wooden ships from the Roman Empire which were besieging Syracuse. (Although no proof of such a feat exists, the Greek navy recreated the experiment in 1973 and successfully set fire to a wooden boat at a distance of 50 meters.)

20 A.D.
Chinese document use of burning mirrors to light torches for religious purposes.

1st to 4th Century A.D.
The famous Roman bathhouses in the first to fourth centuries A.D. had large south facing windows to let in the sun's warmth. For an example, see information on the <http://www.hum.huji.ac.il/archaeology/zippori/RomanSeph.htm> Zippori in the Roman Period from the Hebrew University of Jerusalem.

6th Century A.D.
Sunrooms on houses and public buildings were so common that the Justinian Code initiated "sun rights" to ensure individual access to the sun.

1200s A.D.
Ancestors of Pueblo people called Anasazi in North America live in south-facing cliff dwellings that capture the winter sun.

William Grylls Adams,
Courtesy of John Perlin
2002 From Space to Earth:
The Story of Solar Electricity

Heinrich Hertz,
Courtesy of NASA/
Goddard Space
Flight Center

Solar Water Heater
Courtesy of John Perlin/
Butti Solar Archives

Albert Einstein, courtesy of
the Lotte Jacobi Archives,
University of Hampshire

$$a = \frac{F}{M} \left(1 - \frac{v^2}{c^2}\right)^{3/2}$$

Theory of
Relativity equation

1839
French scientist Edmond Becquerel discovers the photovoltaic effect while experimenting with an electrolytic cell made up of two metal electrodes placed in an electricity-conducting solution—electricity-generation increased when exposed to light.

1860s
French mathematician August Mouchet proposed an idea for solar-powered steam engines. In the following two decades, he and his assistant, Abel Pifre, constructed the first solar powered engines and used them for a variety of applications. These engines became the predecessors of modern parabolic dish collectors.

1873
Willoughby Smith discovered the photoconductivity of selenium.

1876
1876 William Grylls Adams and Richard Evans Day discover that selenium produces electricity when exposed to light. Although selenium solar cells failed to convert enough sunlight to power electrical equipment, they proved that a solid material could change light into electricity without heat or moving parts.

1883
Charles Fritts, an American inventor, described the first solar cells made from selenium wafers.

1891
Baltimore inventor Clarence Kemp patented the first commercial solar water heater. For more information on the water heater, see the http://www.californiasolarcenter.org/history_solarthermal.html California Solar Center.

1905
Albert Einstein published his paper on the photoelectric effect (along with a paper on his theory of relativity).

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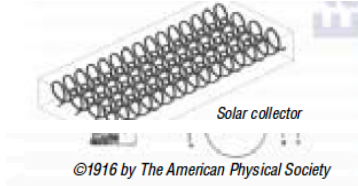
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1908

1908 William J. Bailey of the Carnegie Steel Company invents a solar collector with copper coils and an insulated box—roughly, it's present design.

1918

Polish scientist Jan Czochralski developed a way to grow single-crystal silicon. For more information on Czochralski, see the article <http://rekt.pol.lublin.pl/users/ptwk/art2.htm> Professor Jan Czochralski (1885-1953) and His Contribution to the Art and Science of Crystal Growth.



Jan Czochralski, courtesy of Debra Kaiser, AAGC newsletter

1921

Albert Einstein wins the Nobel Prize for his theories (1904 research and technical paper) explaining the photoelectric effect.

1932

Audobert and Stora discover the photovoltaic effect in cadmium sulfide (CdS).

1954

1954 Photovoltaic technology is born in the United States when Daryl Chapin, Calvin Fuller, and Gerald Pearson develop the silicon photovoltaic (PV) cell at Bell Labs—the first solar cell capable of converting enough of the sun's energy into power to run everyday electrical equipment. Bell Telephone Laboratories produced a silicon solar cell with 4% efficiency and later achieved 11% efficiency. See the http://www.californiasolarcenter.org/history_pv.html for more information.



Bell Labs scientists, Daryl Chapin, Calvin Fuller, and Gerald Pearson. courtesy of

1963

Sharp Corporation succeeds in producing practical silicon photovoltaic modules.



1974-1975

1976

David Carlson and Christopher Wronski, RCA Laboratories, fabricate first amorphous silicon photovoltaic cells.



arren Gretz, NREL / PIX04501

EPFL Photovoltaic effect

The **photovoltaic effect** is a process that generates voltage or electric current in a photovoltaic cell when it is exposed to sunlight. It is this effect that makes solar panels useful, as it is how the cells within the panel convert sunlight to electrical energy. The photovoltaic effect was first discovered in 1839 by Edmond Becquerel.

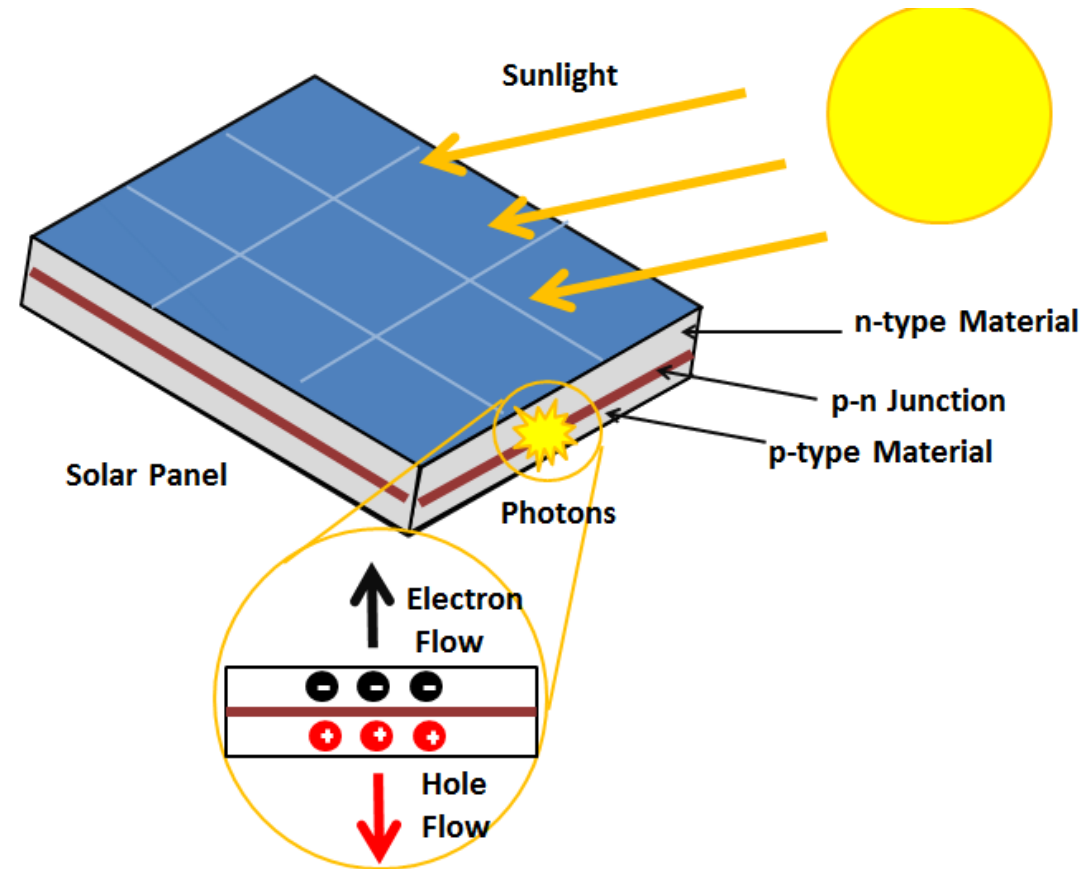
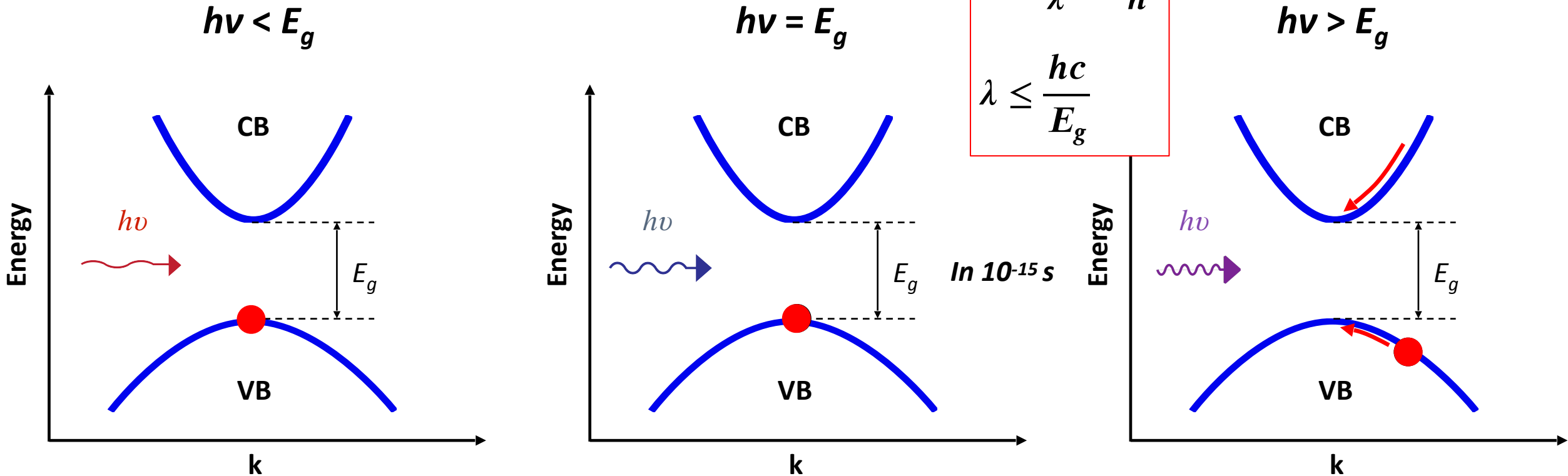


Image taken from https://energyeducation.ca/encyclopedia/Photovoltaic_effect

Photoelectric effect is the emission of electrons from the surface of a substance in response to incident light	Photovoltaic effect is the process (e.g. in pn junction) produce an electrical voltage under illumination
Electrons are emitted	Electrons are not emitted
An electric current is not generated	An electric current is generated
Occurs when the energy provided by photons is enough to overcome the electron binding energy (W_F)	Occurs when the energy provided by photons is enough to overcome the potential barrier of excitation (band gap)

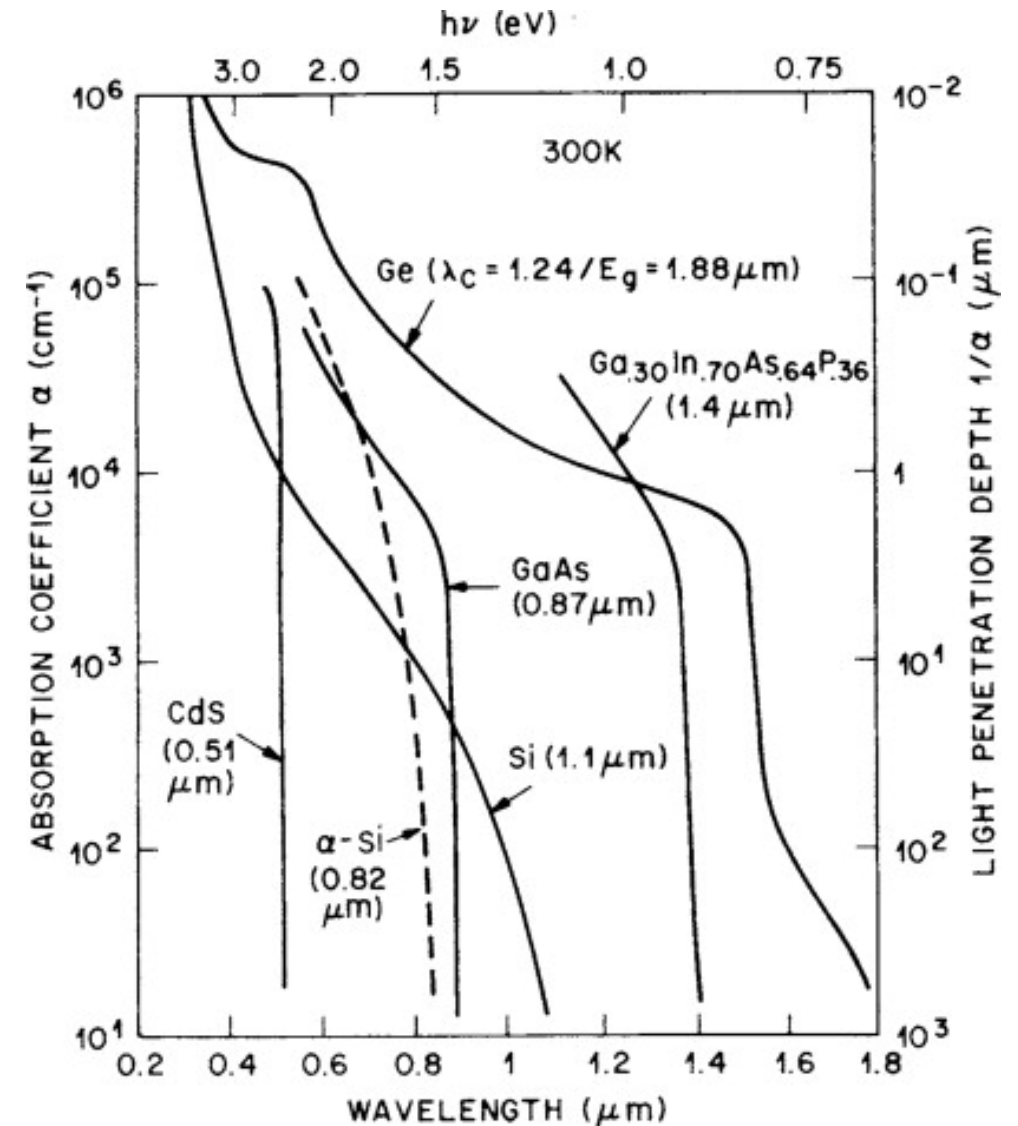
$$\nu = \frac{c}{\lambda} \geq \frac{E_g}{h}$$

$$\lambda \leq \frac{hc}{E_g}$$



- Photons are absorbed by the SC and create electron-hole pairs when the photon energy is equal and greater than the band gap of the SC.
- If the photon energy is much greater than the band gap, the excess energy ($h\nu - E_g$) is lost due to the ultrafast relaxation of carriers, e.g. 10^{-14} to 10^{-11} sec (thermalization or vibrational relaxation).

- **Absorption** of photons with energy $\geq E_g$
- Loss: **Transmission, Reflection**
- When an object is exposed to radiation, some of the incident radiation is absorbed (**A**), some are scattered (**R**), and some are transmitted, Transmittance (**T**). (**A + T + R = 1**)
- Absorption coefficient α : Absorption coefficient characterizes the efficiency of a material in absorbing optical power.
- Absorption depth is the inverse of absorption coefficient, describing how deeply light penetrates into a SC.
- Direct gap transitions are much more efficient than indirect transitions and results in much higher α .



- **Lambert's law:** the decrease in light intensity due to the thickness of the absorbing medium at any point is directly proportional to the light intensity.

$$-\frac{dI}{dl} \propto I$$

dI : the slight decrease in light intensity when passing a small distance dl

I : the intensity of monochromatic light just before entering the medium

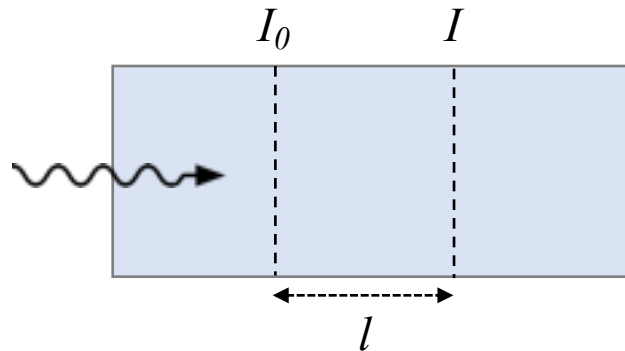
$$-\frac{dI}{dl} = \alpha I$$

α : absorption coefficient (cm^{-1})

$$-\ln I = \alpha l + C$$

At $l = 0, I = I_0 \rightarrow C = -\ln I_0$

$$-\ln\left(\frac{I}{I_0}\right) = \alpha l \quad \rightarrow \quad I = I_0 \exp(-\alpha l) \quad \text{or} \quad \log\left(\frac{I}{I_0}\right) = -\alpha' l \quad \text{where } \alpha' : \text{extinction coefficient} = \alpha/2.303$$



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- **Beer-Lambert's law:** Extended by Beer, the decrease in light intensity due to not only the thickness of the absorbing medium but also the concentration of the solution.

$$-\frac{dI}{dl} \propto c$$

\rightarrow

$$\log\left(\frac{I}{I_0}\right) = -\epsilon c l$$

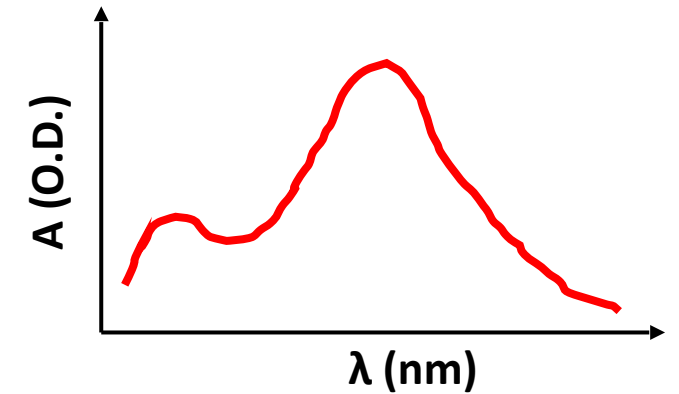
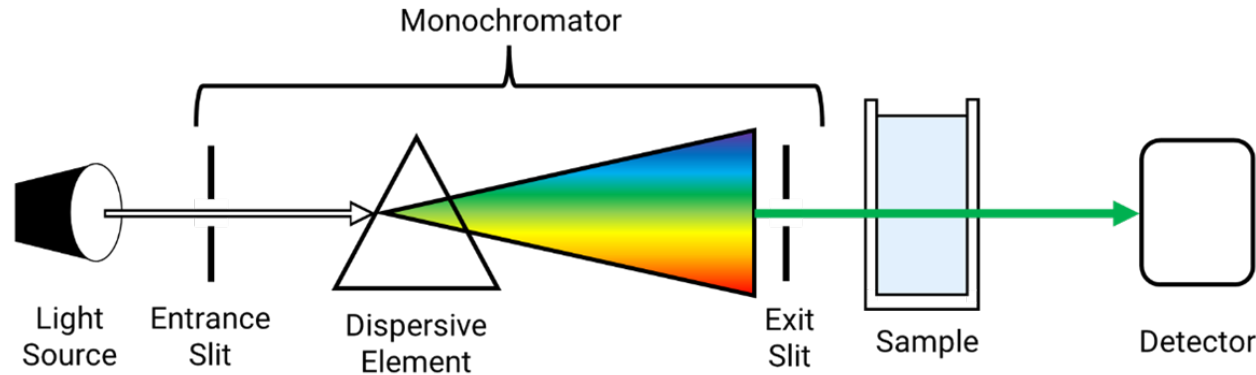
ϵ = molar extinction coefficient ($\text{Lmol}^{-1}\text{cm}^{-1}$)

c = concentration (M)

l is the path length (cm)

Molar extinction coefficient: It measures the probability of electronic transitions.

- Absorbance (**A**) known as optical density (OD): the quantity of light absorbed by a sample.
- Transmittance (**T**): the quantity of light that passes through a sample.
- Tool: UV-vis-NIR spectroscopy



$$T = \frac{I}{I_0} \quad \text{or} \quad T(\%) = 100 \times T$$

$$A = -\log\left(\frac{I}{I_0}\right) = -\log(T) = \log\left(\frac{1}{T}\right)$$

Beer-Lambert's law

From absorbance definition

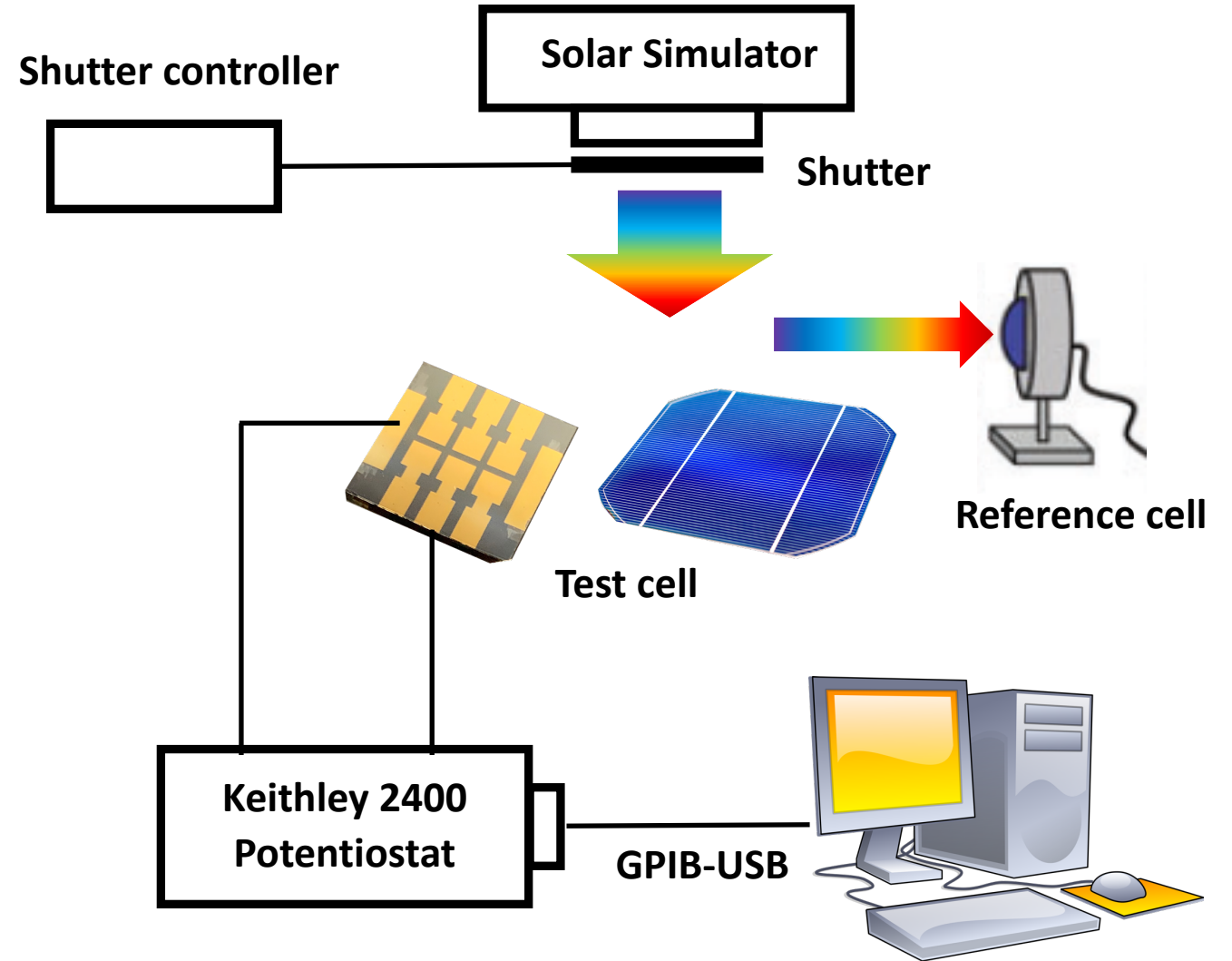
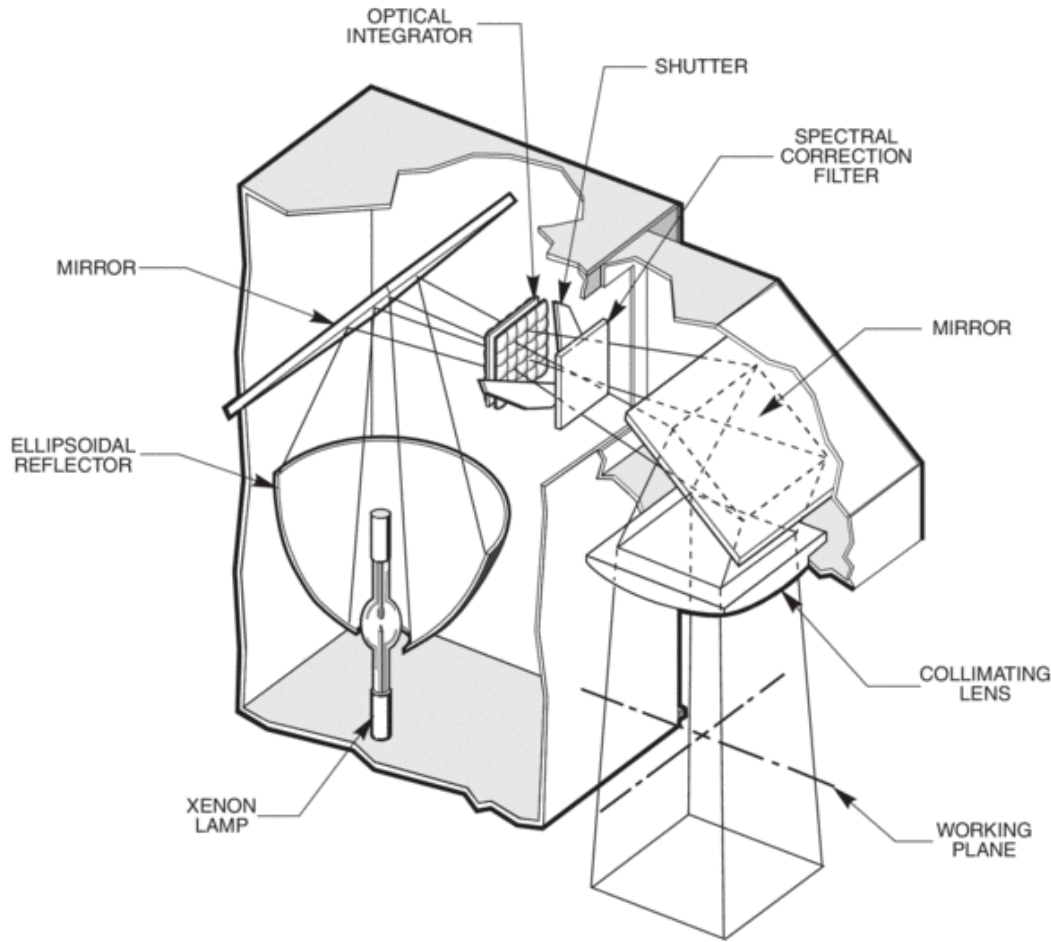
The relative loss of intensity (**absorptance**): $1 - T = 1 - \frac{I}{I_0} = 1 - 10^{-A}$ $A = -\log\left(\frac{I}{I_0}\right)$

Example 1) Find the relative amount of light that gets absorbed by the sample (**absorptance**) if the absorbance of the sample is **2** at a particular wavelength.

$$1 - 10^{-A} = 1 - \frac{1}{100} = 0.99 \quad (\text{with } A = 2)$$

99% of the light is absorbed and 1% of light is transmitted.

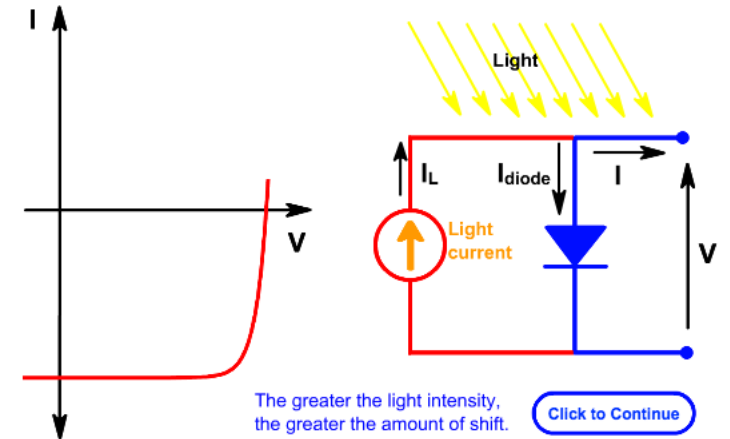
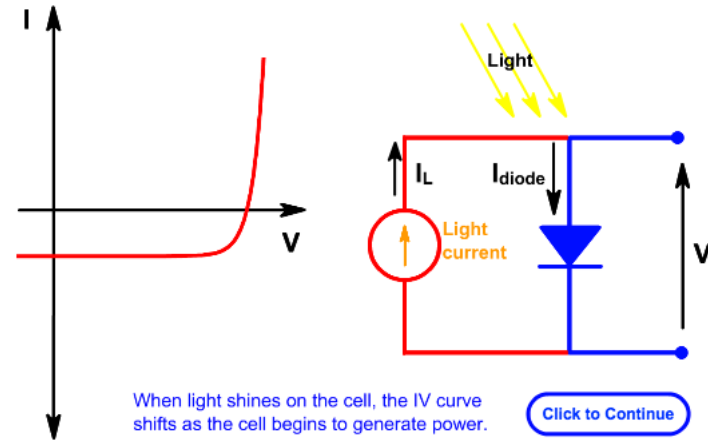
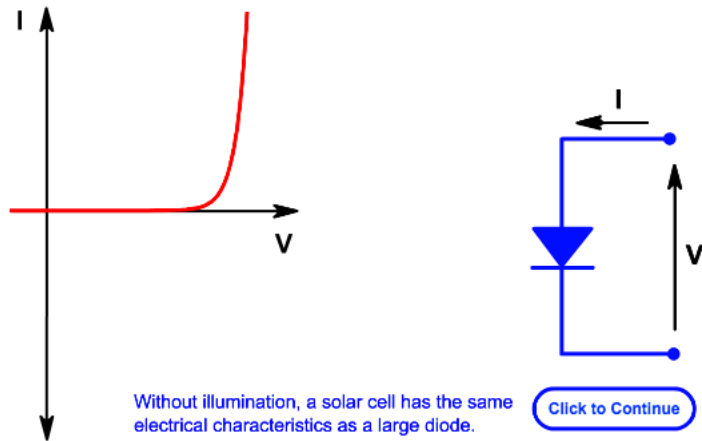
Absorbance (OD)	Transmittance (%)	(Absorptance)
0	100%	0%
1	10%	90%
2	1%	99%
3	0.1%	99.9%
5	0.001%	99.999%



The IV curve of a solar cell = the superposition of the IV curve of the solar cell diode in the dark with the light-generated current.

$$I = I_L - I_D(V) = I_L - I_0 \left[\exp\left(\frac{eV}{nk_B T}\right) - 1 \right]$$

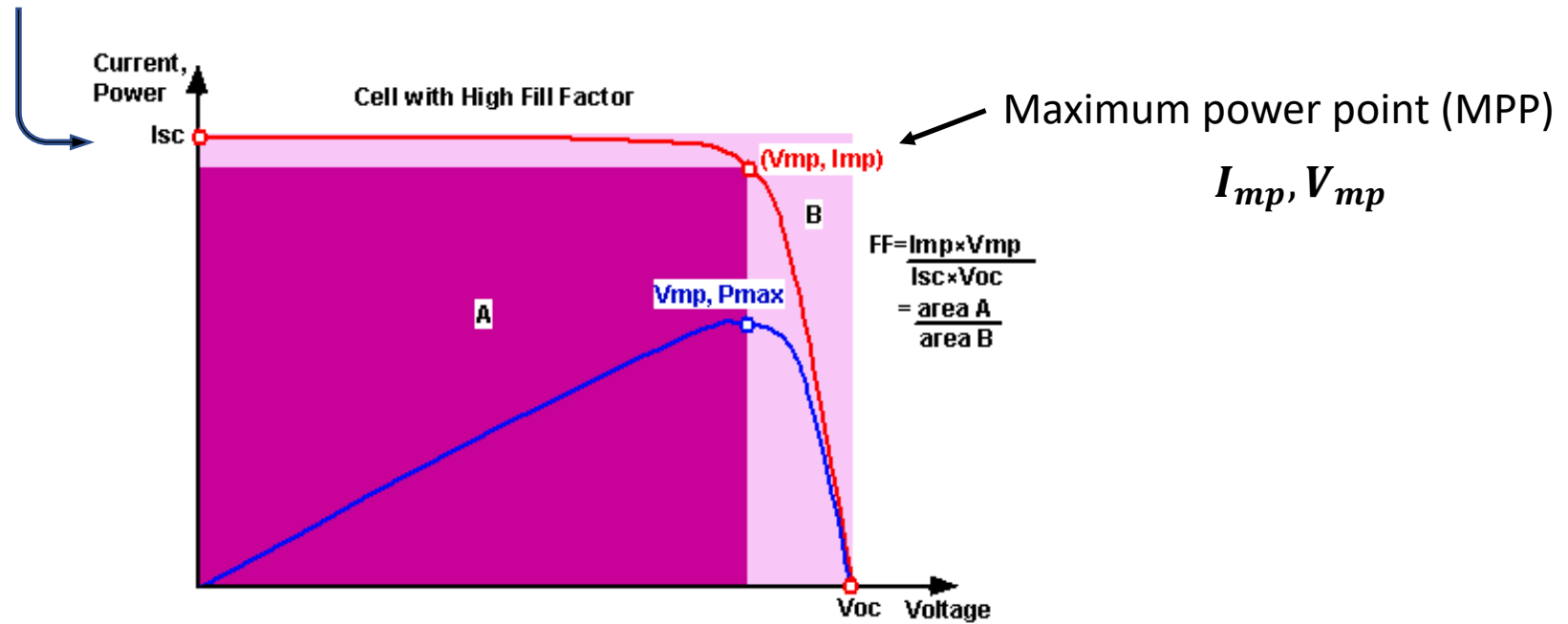
I_L = Light generated current
 I_D = Dark current



From PV education

EPFL IV curve: I_{sc} V_{oc} and MPP

At $V = 0$, the condition is defined as the **short-circuit condition: short-circuit current (I_{sc})**



From PV education

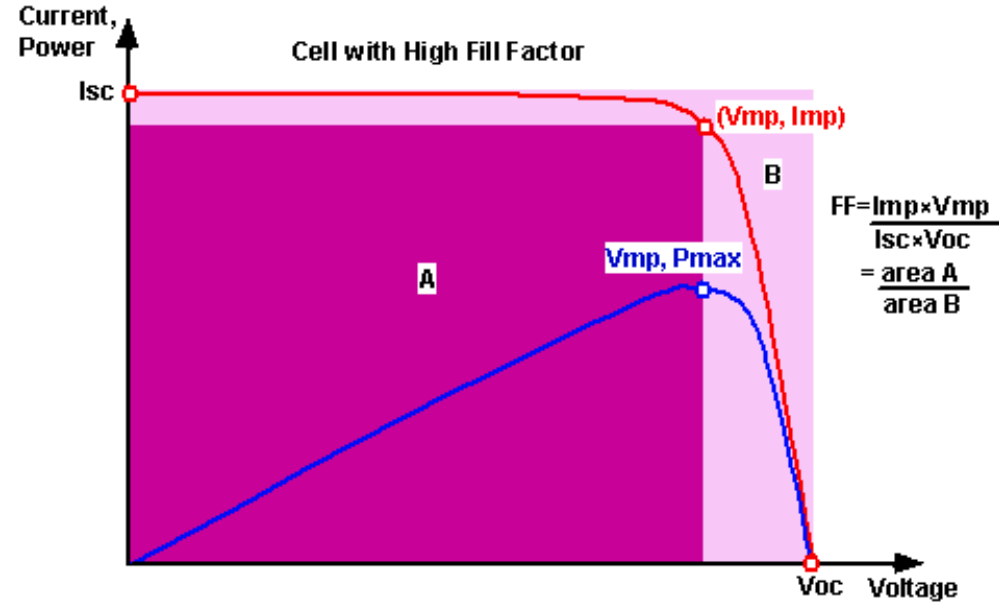
At $I = 0$, the condition is defined as the **open-circuit condition: open-circuit voltage (V_{oc})**

$$V_{oc} = \frac{nk_B T}{e} \ln\left(\frac{I_L}{I_0} + 1\right)$$

IV curve: Fill Factor and PCE

Fill Factor (FF) is a parameter which determines the maximum power from a solar cell.

$$FF = \frac{P_{mp}}{V_{OC} \times I_{SC}} = \frac{V_{mp} \times I_{mp}}{V_{OC} \times I_{SC}}$$



- Terrestrial solar cells are measured under **AM1.5G** conditions (**1000 W/m²**) and at a temperature of **25°C**.
- Solar cells intended for space use are measured under AM0 conditions (1367 W/m²).
- A solar cell with a conversion efficiency of 100% is theoretically impossible.

From PV education

$$\text{Power conversion efficiency (\%)} = \eta = \frac{\text{Generated electrical power (W)}}{\text{Incident light power (W)}} \times 100$$

$$\frac{P_{mp}}{P_{inc}} = \frac{V_{OC} \times I_{SC} \times FF}{P_{inc} (W)} = \frac{V_{OC} \times J_{SC} \times FF}{P_{inc} (W/m^2)} \quad \text{with } J_{SC} = I_{SC}/A$$

Intrinsic losses

Unavoidable

- Optical loss
- Transmission loss
- Thermalization loss
- Carnot loss
- Emission loss
- Boltzmann loss

Extrinsic losses

Avoidable

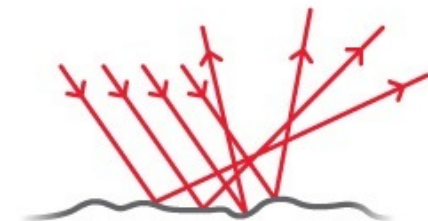
- Non-radiative loss
- Series resistance loss
- Shunt resistance loss
- Parasitic absorption loss

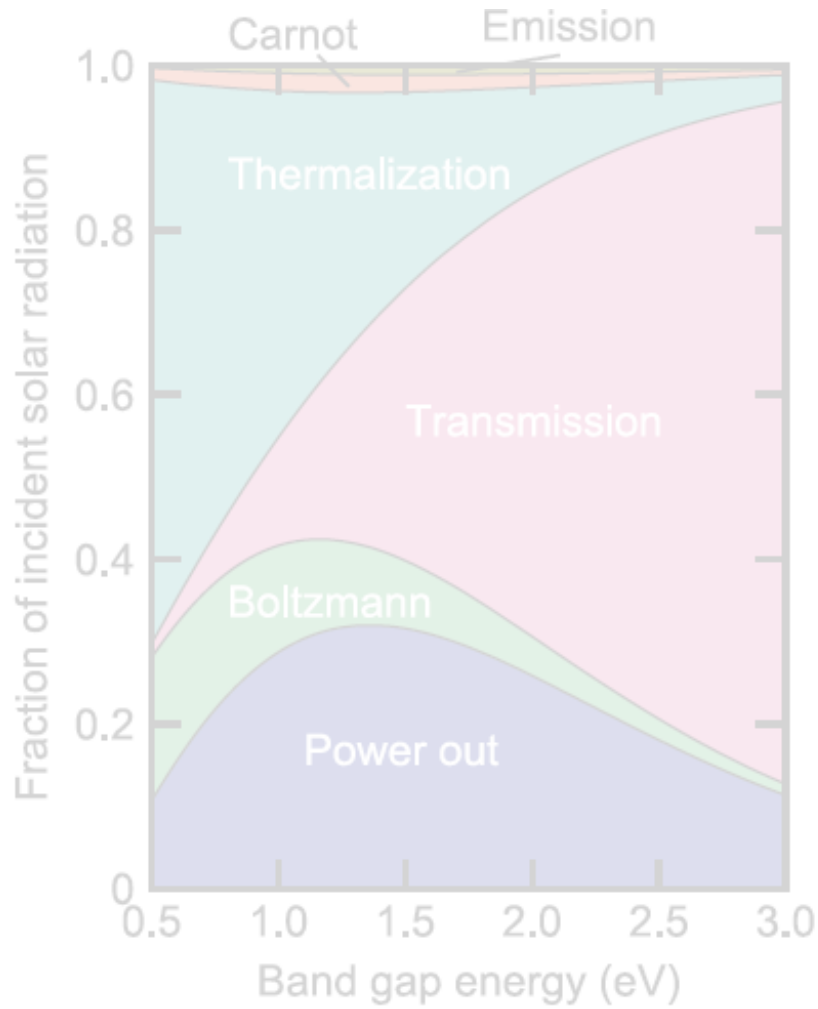
- The optical loss by reflection from the cells.
- The transmission loss by the non-absorption of photons with energy below the bandgap.
- A strong interaction between excited carriers and lattice phonons introduces a thermalization loss as carriers cool to the bandgap edge.

Specular reflection

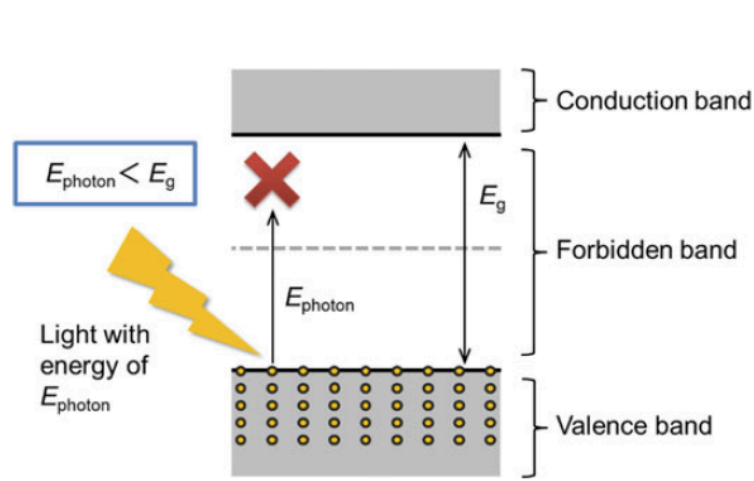


Diffuse reflection

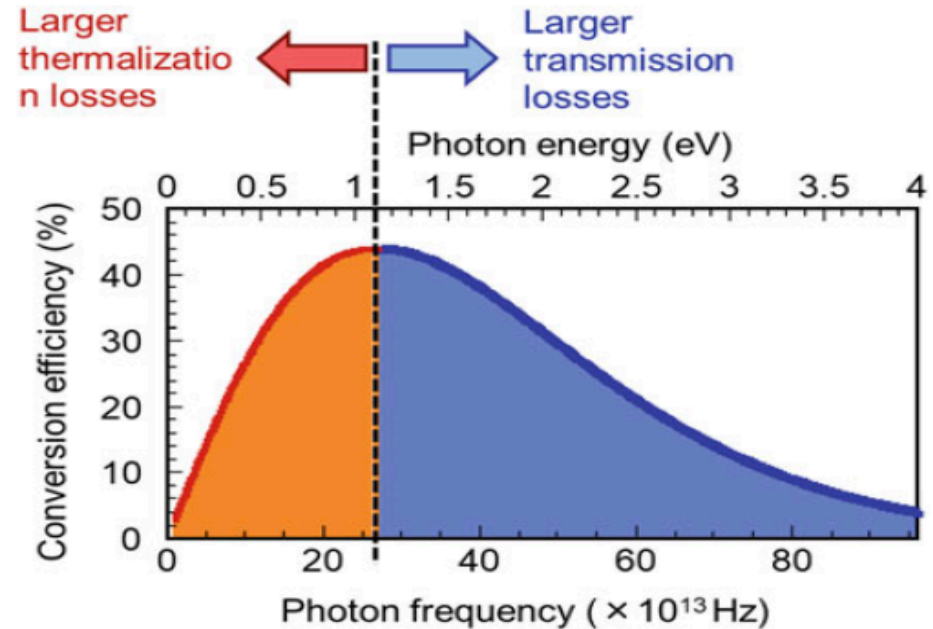
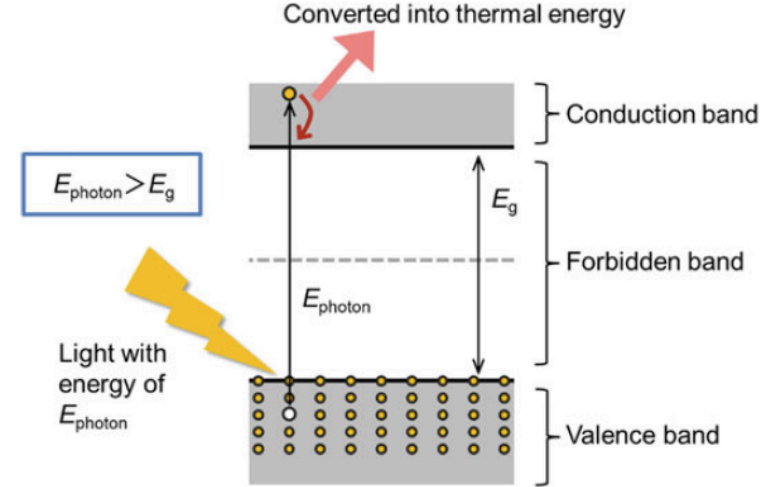




Transmission loss



Thermalization loss



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Extrinsic losses

Avoidable

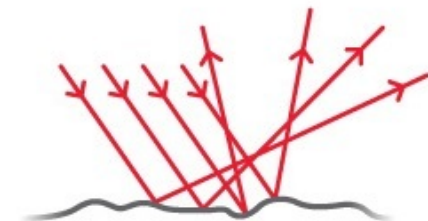
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- The optical loss by reflection from the cells.
- The transmission loss by the non-absorption of photons with energy below the bandgap.
- A strong interaction between excited carriers and lattice phonons introduces a thermalization loss as carriers cool to the bandgap edge.
- Work loss in the heat engine with heat flowing from a hot reservoir (the Sun) to a cold reservoir (the atmosphere).
- The emission loss originates from the photon emission of the cells as a result of radiative recombination.
- Inequality of absorption and emission angles results in an entropy generation.

Specular reflection



Diffuse reflection



$$\eta_c = \frac{W}{Q_1} = \frac{Q_1 - Q_2}{Q_1} = 1 - \frac{Q_2}{Q_1}$$

For a reversible engine, the total entropy is conserved, $S = S_1 - S_2 = 0$,

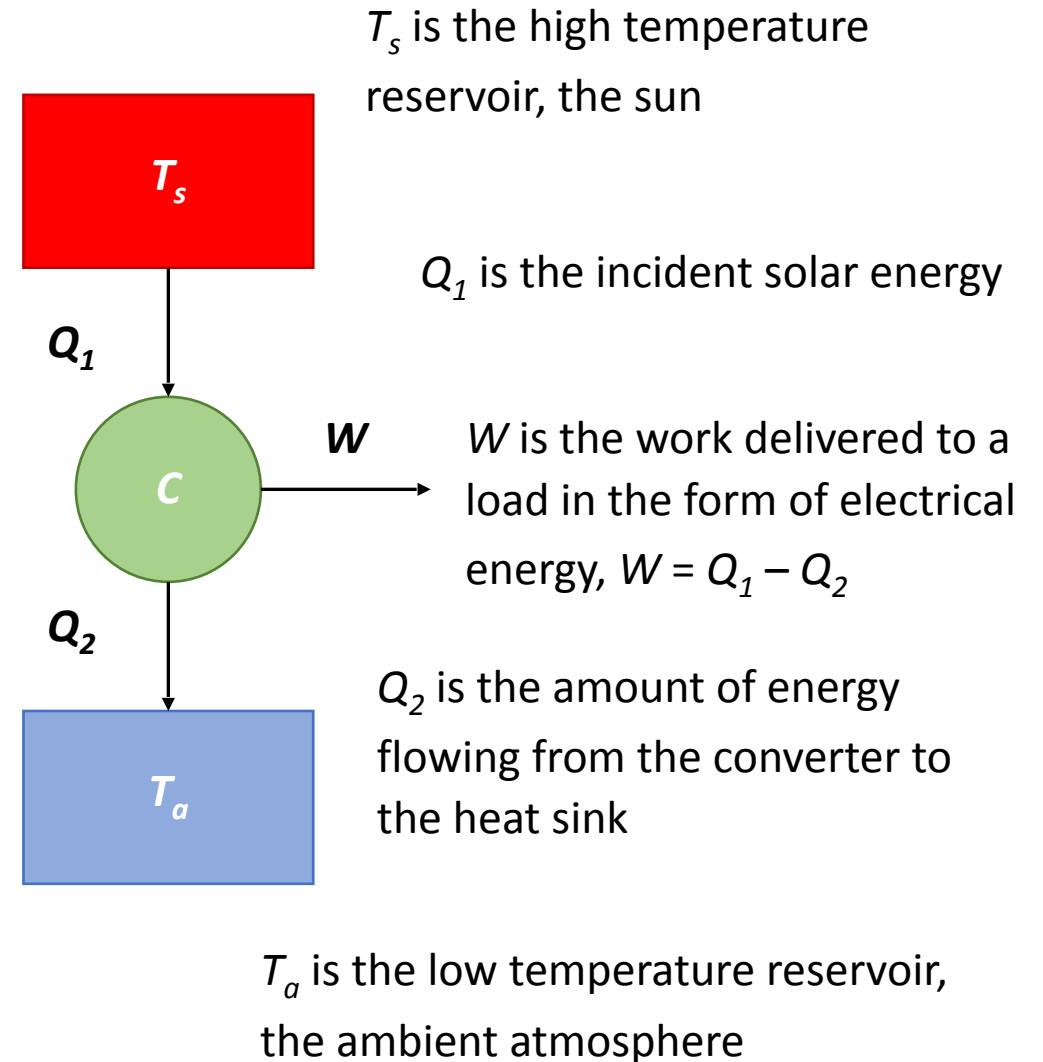
$$\frac{Q_1}{T_1} - \frac{Q_2}{T_2} = 0$$

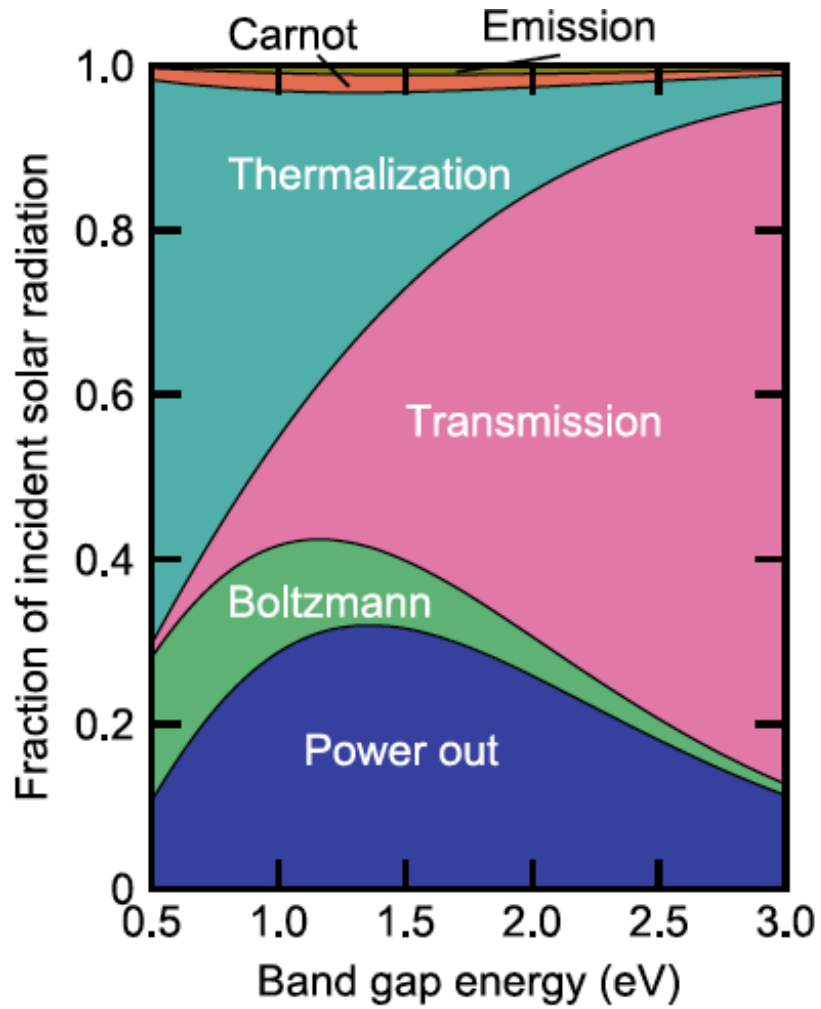
$$\eta_c = 1 - \frac{T_2}{T_1} = 1 - \frac{T_a}{T_s}$$

$T_s = 6,000 \text{ K}$ and $T_a = 300 \text{ K}$

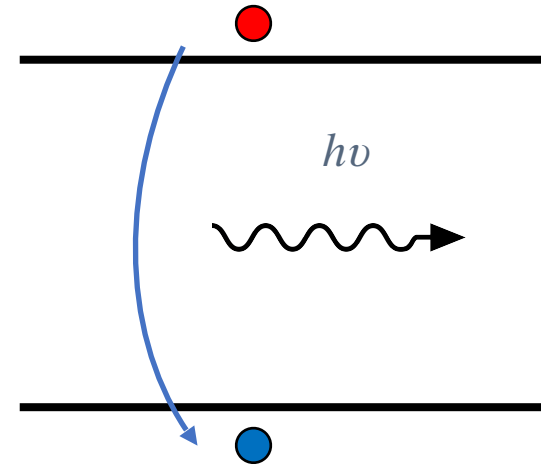
$\eta_c = 0.95$, an upper limit of solar converters

Carnot loss

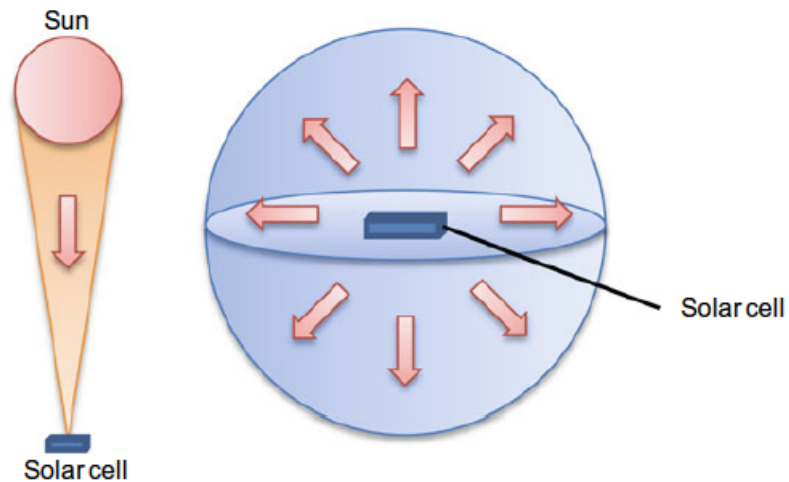




Emission loss



Boltzmann loss: Incident solar energy with the solid angle but the solar cell emits light into the emission solid angle of 2π .



$$\Delta S = k \ln(2\pi) - k \ln(\Omega_{se})$$

Intrinsic losses

Unavoidable

- Optical loss
- Transmission loss
- Thermalization loss
- Carnot loss
- Emission loss
- Boltzmann loss

Extrinsic losses

Avoidable

- Non-radiative loss
- Series resistance loss
- Shunt resistance loss
- Parasitic absorption loss

- Non-radiative recombination (NRR) losses originating from different mechanisms: Shockley-Read-Hall (SRH) recombination, Auger recombination, Recombination at Interface, etc.

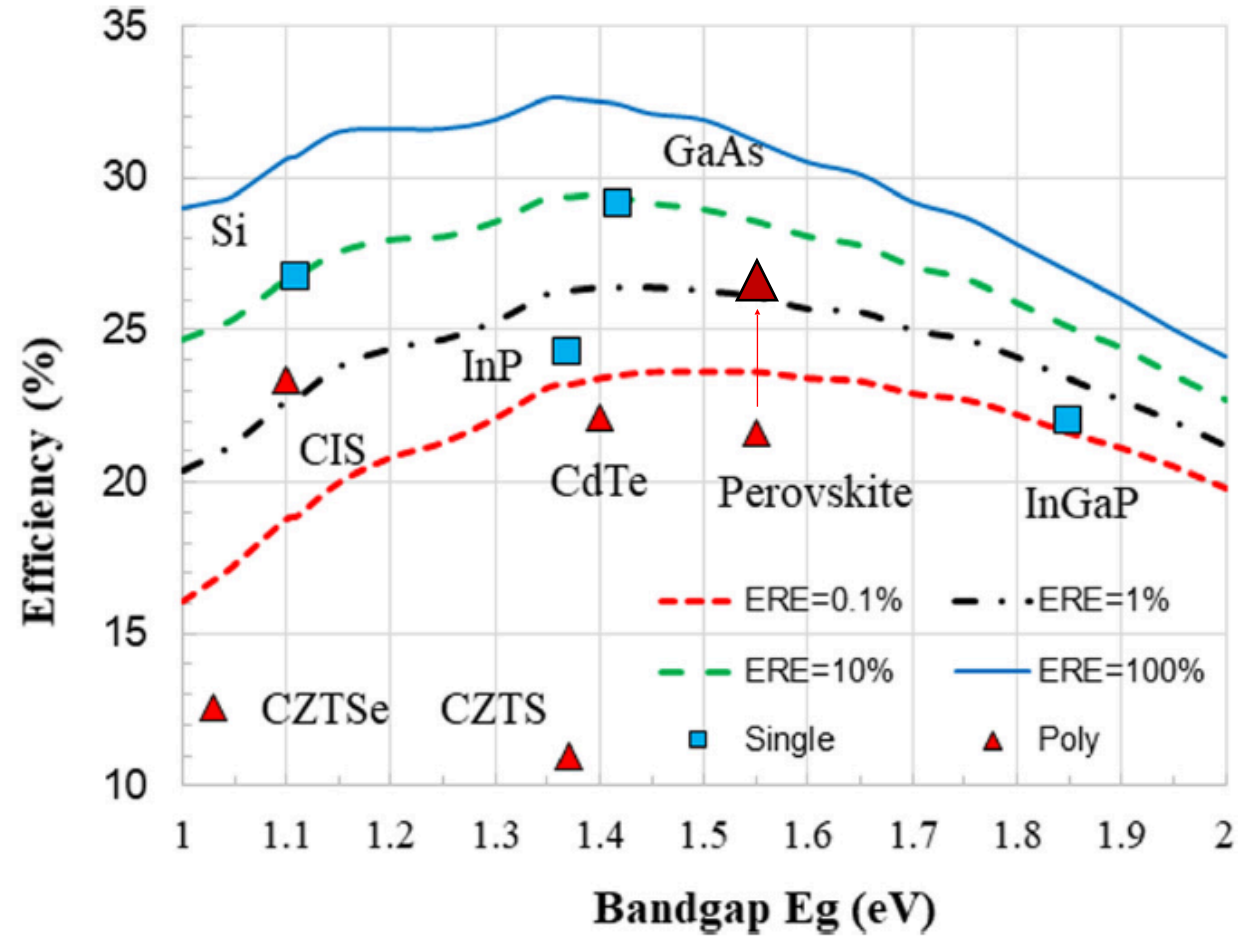
External radiative efficiency (ERE) is the ratio of radiatively recombined carriers against all recombined carriers.

$$ERE = J_{rad} / (J_{rad} + J_{nrad})$$

Therefore the rest fraction of the recombinations ($1 - ERE$) refers to the non-radiative recombinations.

$ERE = 1$ (100%) at Shockley-Queisser limit

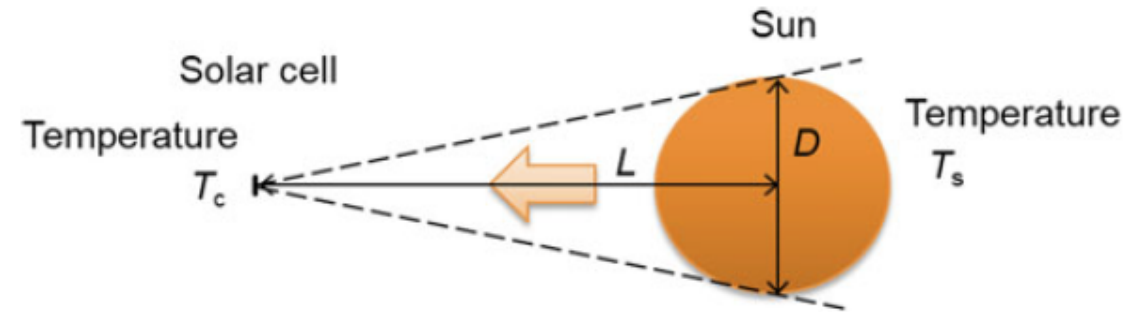
- The series resistance in solar cells.
- The shunt resistance in solar cells.
- Parasitic absorption by solar cell component (non active).



$$ERE = J_{rad} / (J_{rad} + J_{nrad})$$

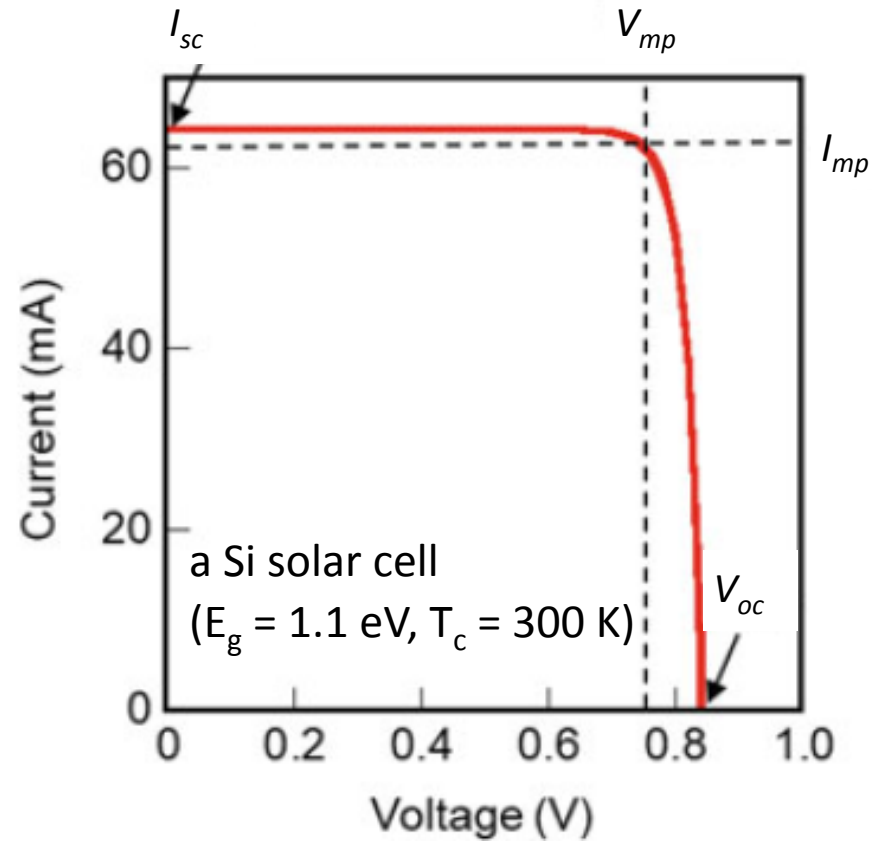
Yamaguchi et al., J. Appl. Phys. 129, 240901 (2021)

The balance between the carrier generation and recombination that occurs under the particular condition (maximum power point) for maximum extractable electrical power was considered.

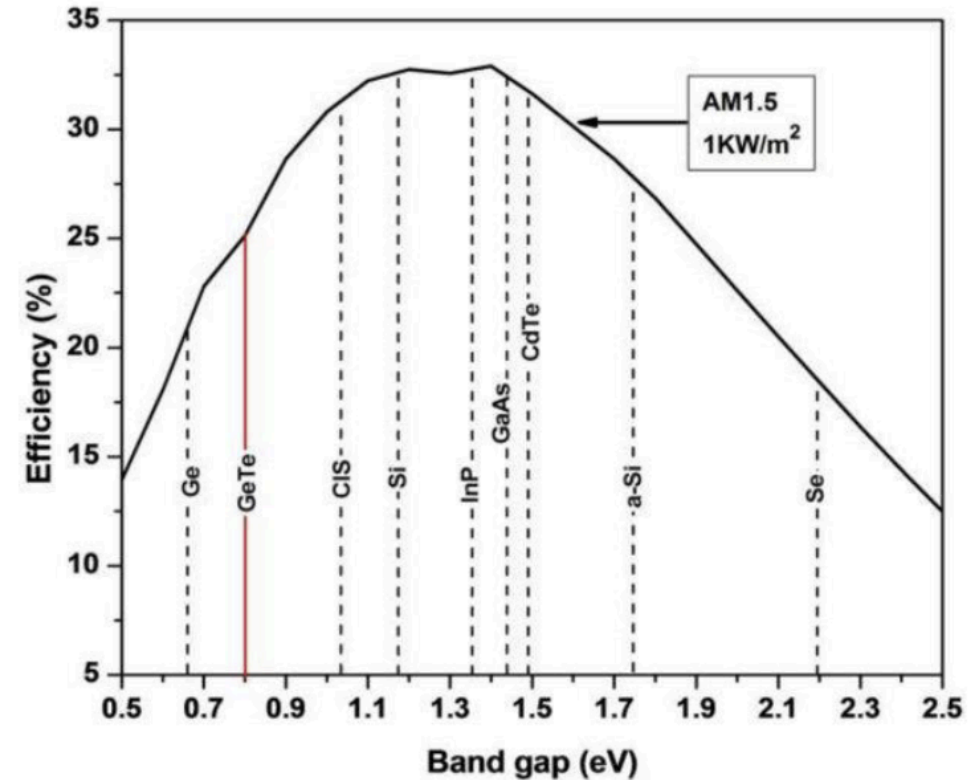


- The solar cell has a finite temperature T_c and also emits black-body radiation = the solar cell is surrounded by a black body with T_c .
- Consider electron-hole pairs generated by the radiation of the black body with T_c .
- The probability t_c for an electron-hole pair by the incident photons larger than E_g is set to 100%, $t_c = 1$.
- All photogenerated electrons and holes thermalize to the band edges.
- All the photogenerated charge carriers are collected at short-circuit condition = unlimited mobility.
- The radiation absorbed by the cell is equal to the radiation emitted ($R_{\text{abs}} = R_{\text{emit}}$) = non-radiative recombination is not taken into account.

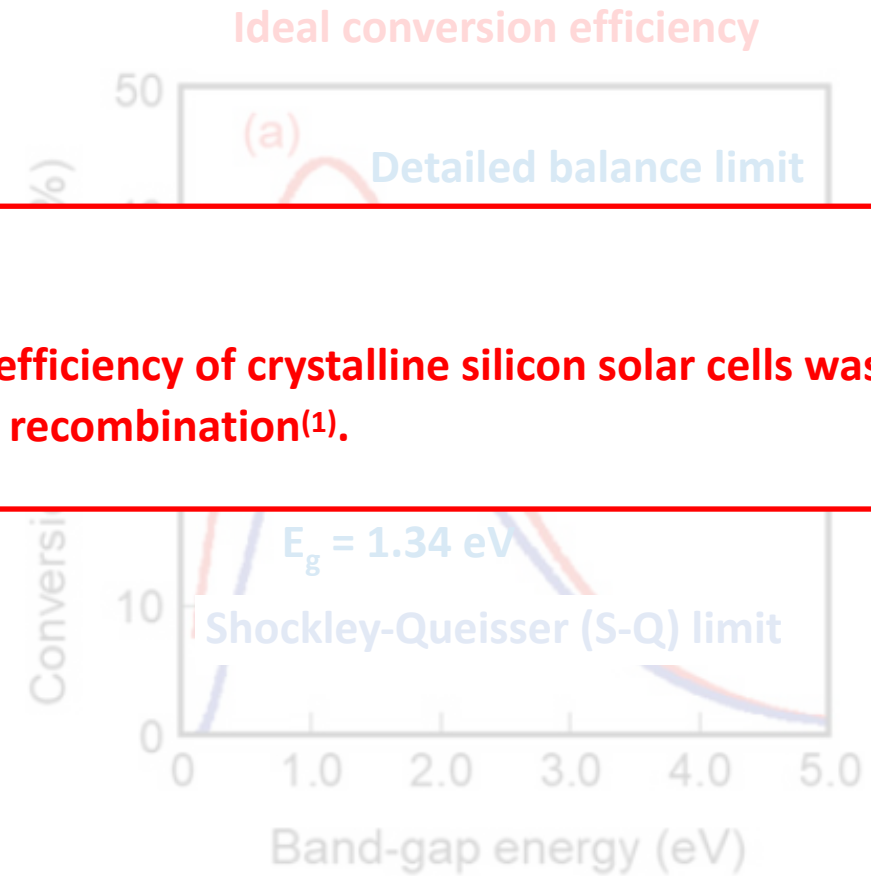
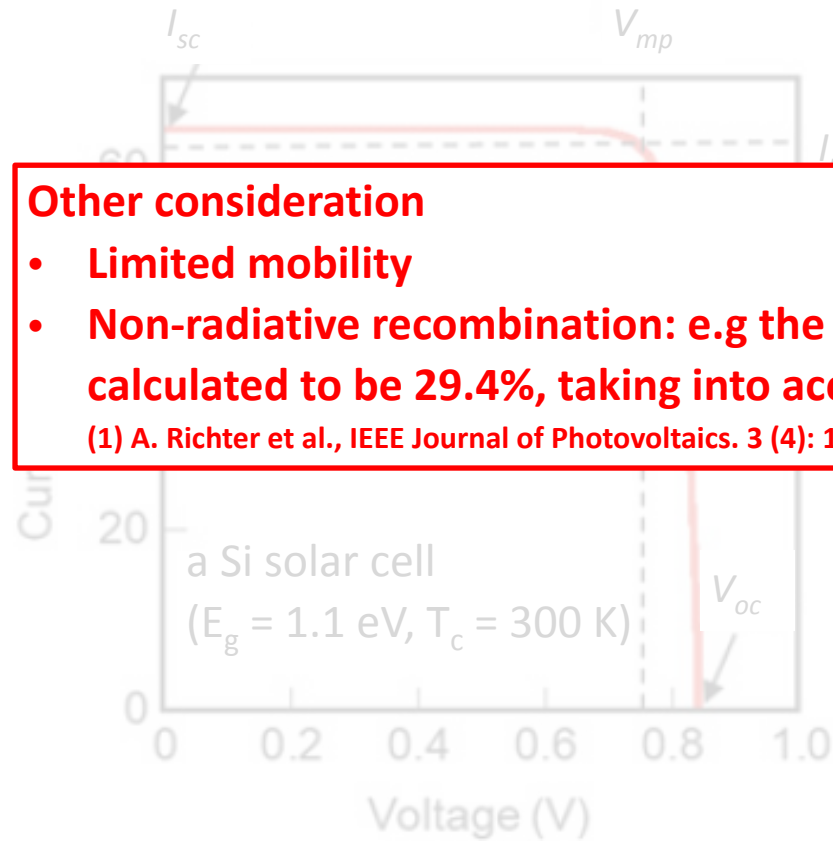
$$IV = I_0V \left[\exp\left(\frac{V_{oc}}{V_c}\right) - \exp\left(\frac{V}{V_c}\right) \right]$$



Shockley-Queisser (S-Q) limit

33.7% with $E_g = 1.34$ eV32% with Si $E_g = 1.12$ eV

$$IV = I_0V \left[\exp\left(\frac{V_{oc}}{V_c}\right) - \exp\left(\frac{V}{V_c}\right) \right]$$



Other consideration

- Limited mobility
- Non-radiative recombination: e.g the theoretical efficiency of crystalline silicon solar cells was calculated to be 29.4%, taking into account Auger recombination⁽¹⁾.

(1) A. Richter et al., IEEE Journal of Photovoltaics. 3 (4): 1184–1191 (2013)

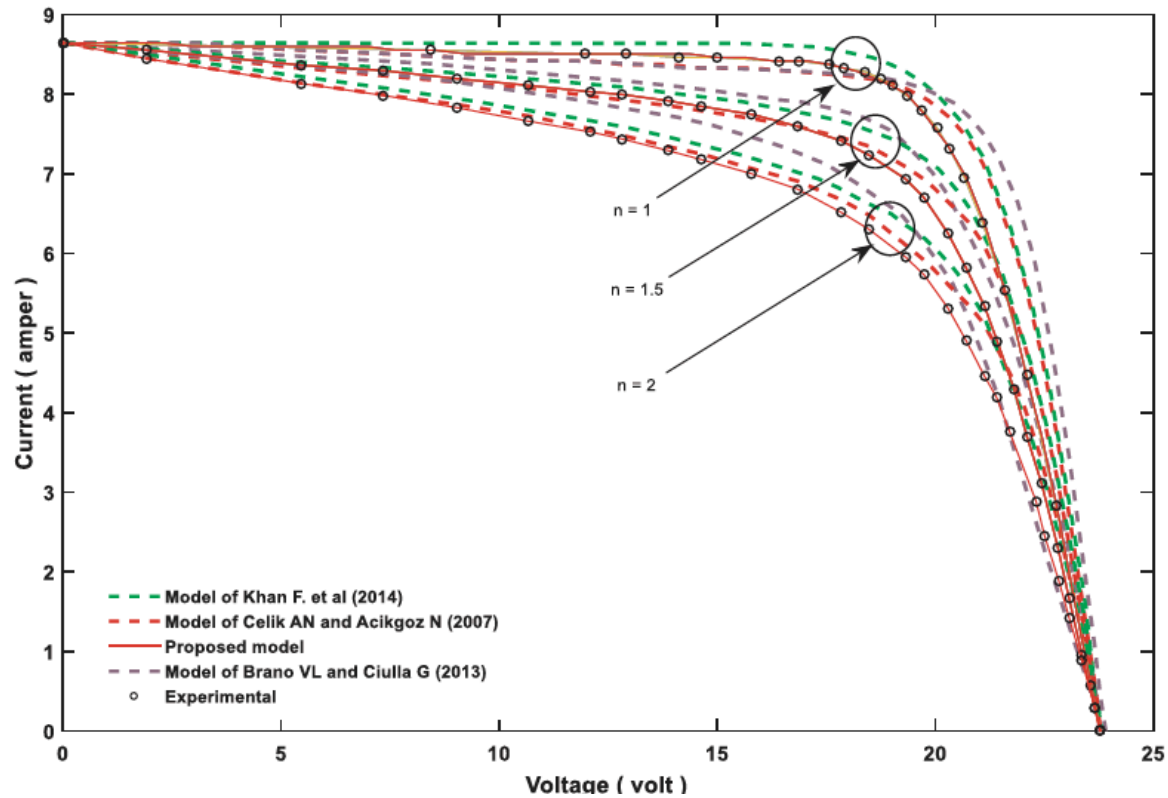
$$I = I_L - I_0 \left[\exp\left(\frac{eV}{nk_B T}\right) - 1 \right] \approx I_L - I_0 \left[\exp\left(\frac{eV}{nk_B T}\right) \right]$$

n = A measure of how closely the diode follows the ideal diode equation.
Ideality factor depends on recombination type.

Recombination type	n	Description
band to band (low level injection)	1	Recombination limited by minority carrier.
SRH, band to band (high level injection)	2	Recombination limited by both carrier type.
Junction (depletion region)	2	Two carriers limit recombination.
Auger	2/3	Two majority and one minority carriers required.

$$I = I_L - I_0 \left[\exp \left(\frac{eV}{nk_B T} \right) - 1 \right] \approx I_L - I_0 \left[\exp \left(\frac{eV}{nk_B T} \right) \right]$$

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 Ideality factor depends on recombination type.



Real cells: Resistance of the contacts and leakage currents through the cell and/or around the sides of the device.

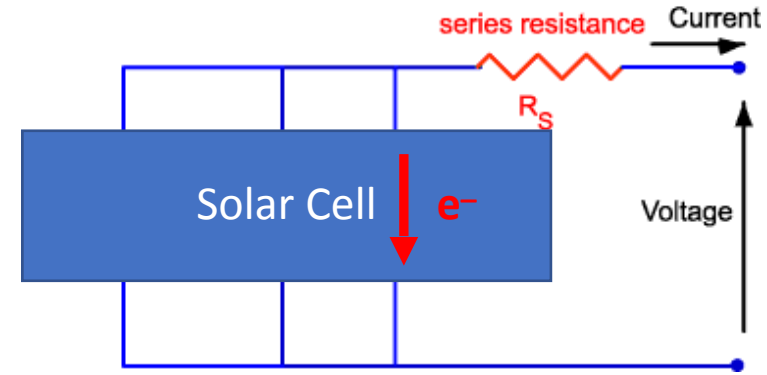
Series resistance (R_{series}): it arises from the resistance of the cell material to current flow, particularly through the front surface to the contacts and from resistive contacts. It is a particular problem at high current densities, i.e. under concentrated light. A **low R_{series}** is desired!

Shunt resistance (R_{shunt}): It arises from the leakage of the current through the cell, e.g. defects or cracks, causing power loss by diverting current. It is a problem with poorly rectifying devices. A **high R_{shunt}** is desired!

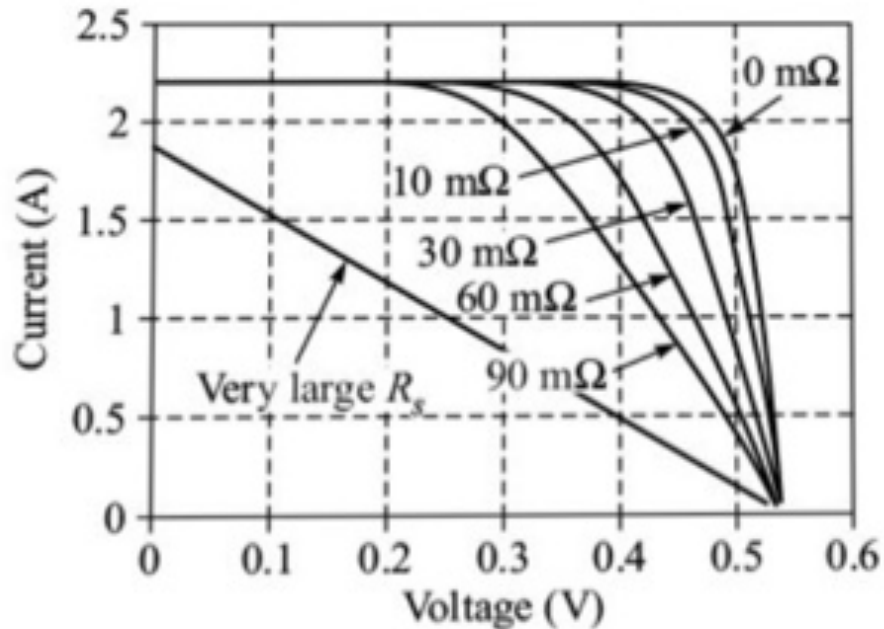
The diode equation becomes:

$$I = I_L - I_0 \left[\exp \left(\frac{e(V + IR_{series})}{nk_B T} \right) - 1 \right] - \frac{V + IR_{series}}{R_{shunt}}$$

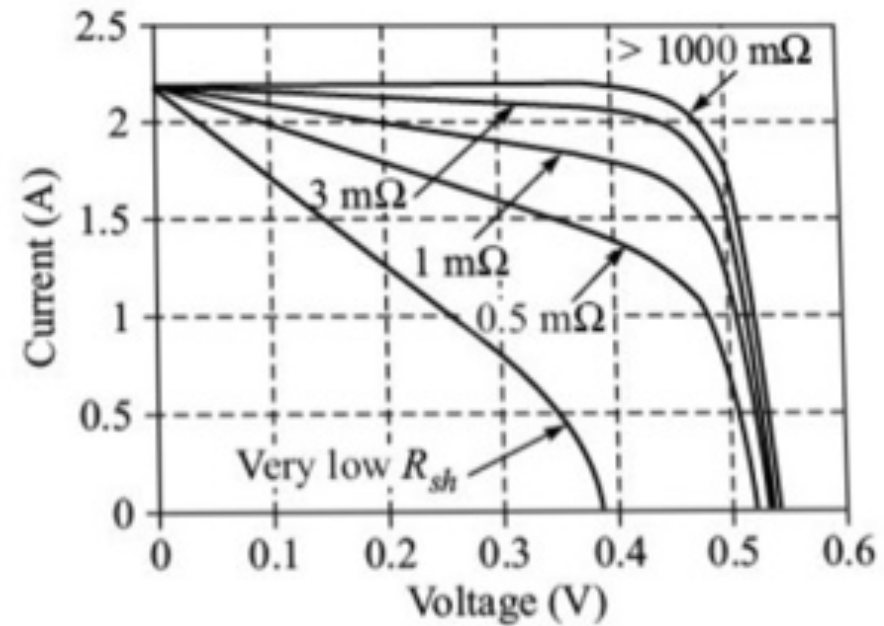
Resistive effects, series resistance (R_{series}) and shunt resistance (R_{shunt}) in solar cells reduce the efficiency of the solar cell by dissipating power in the resistances, mainly the **fill factor**.



The effect of R_{series}

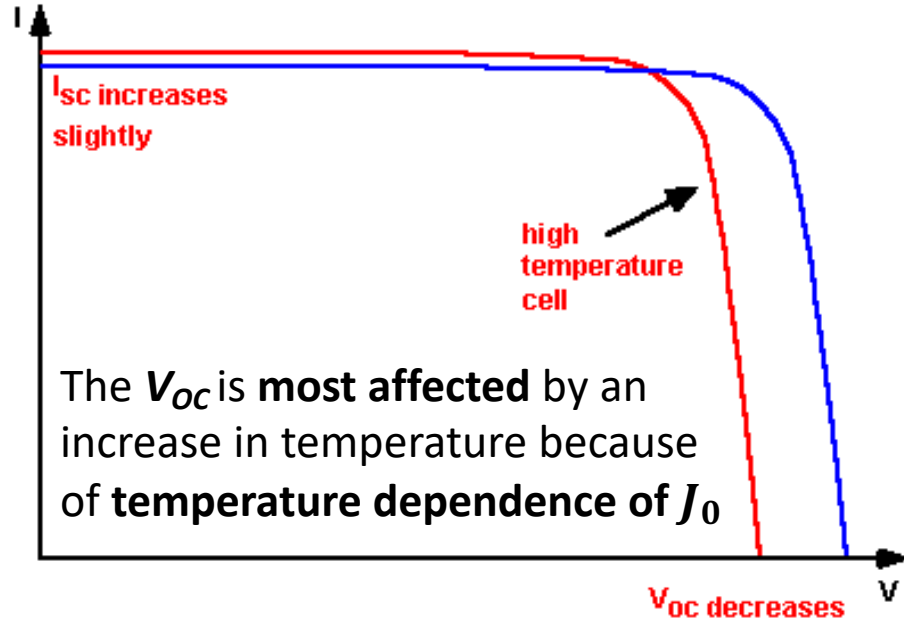


The effect of R_{shunt}



$$\frac{1}{J_{sc}} \frac{dJ_{sc}}{dT} \approx 0.0006 \text{ per } ^\circ\text{C for Si}$$

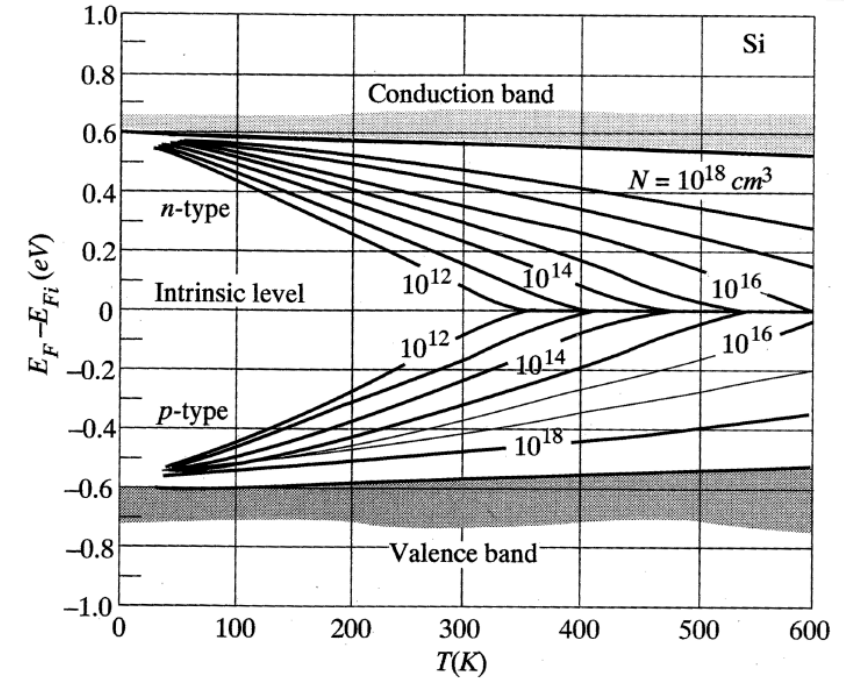
(0.06% per °C)



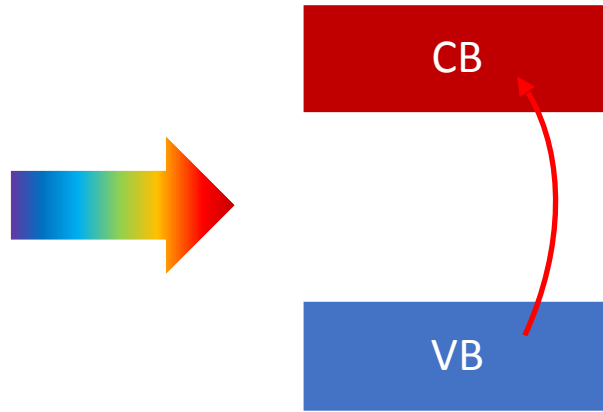
The **V_{oc}** is most affected by an increase in temperature because of temperature dependence of **J₀**

$$\frac{dV_{oc}}{dT} \approx -2.2 \text{ mV per } ^\circ\text{C for Si}$$

An increase in temperature → **Reduced Band Gap of SC**



$$V_{oc} = \frac{nk_B T}{e} \ln\left(\frac{I_L}{I_0} + 1\right) \quad J_0 = e \left(\frac{D_e n_{p0}}{L_e} + \frac{D_h p_{n0}}{L_h} \right) = e n_i^2 \left(\sqrt{\frac{D_e}{\tau_e} \frac{1}{N_A}} + \sqrt{\frac{D_h}{\tau_h} \frac{1}{N_D}} \right)$$

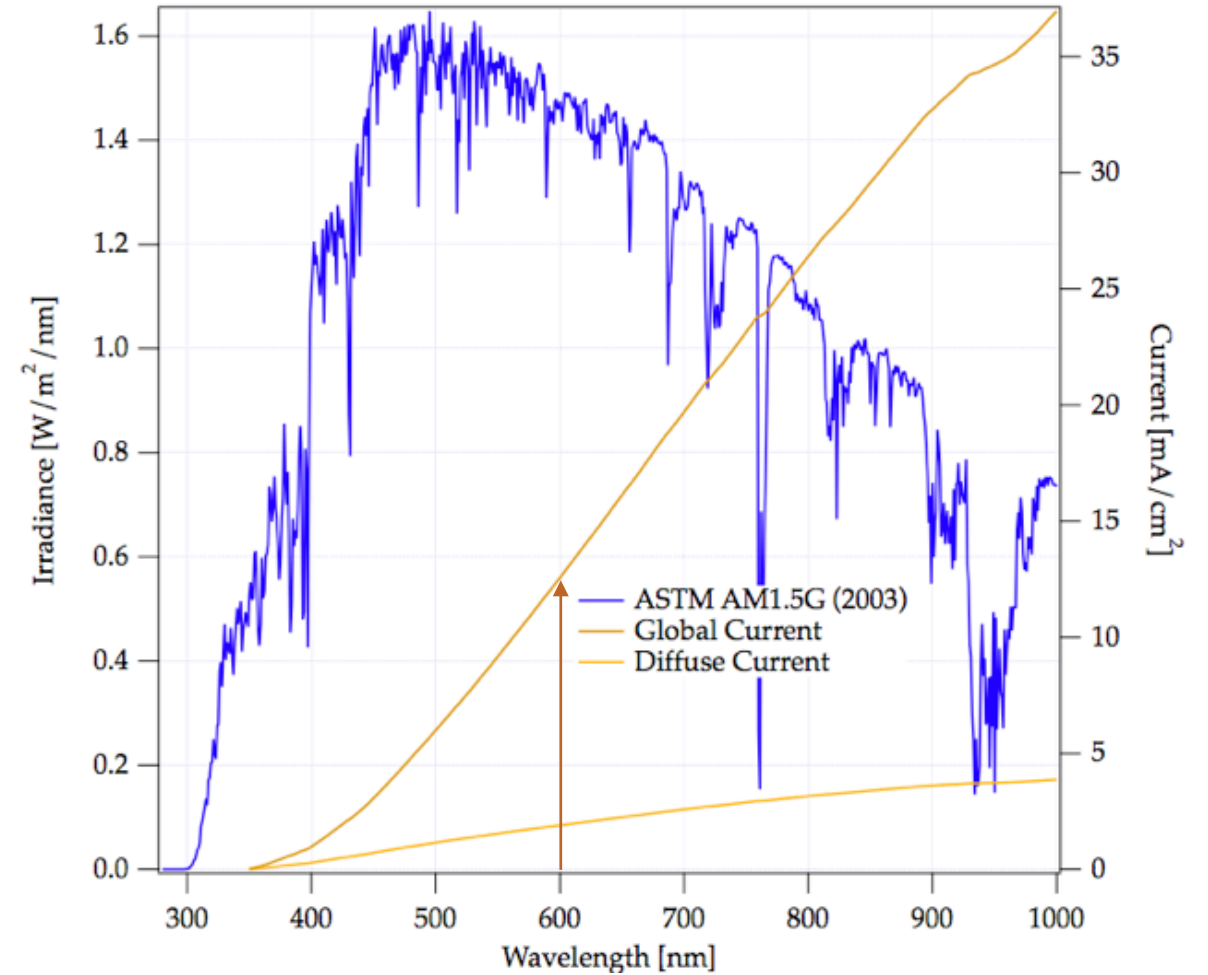


- **Light management:**

- Thickness
- Band gap
- Light management by scattering or reflection

- **Carrier management:**

- Charge separation
- Charge collection



2.1 eV

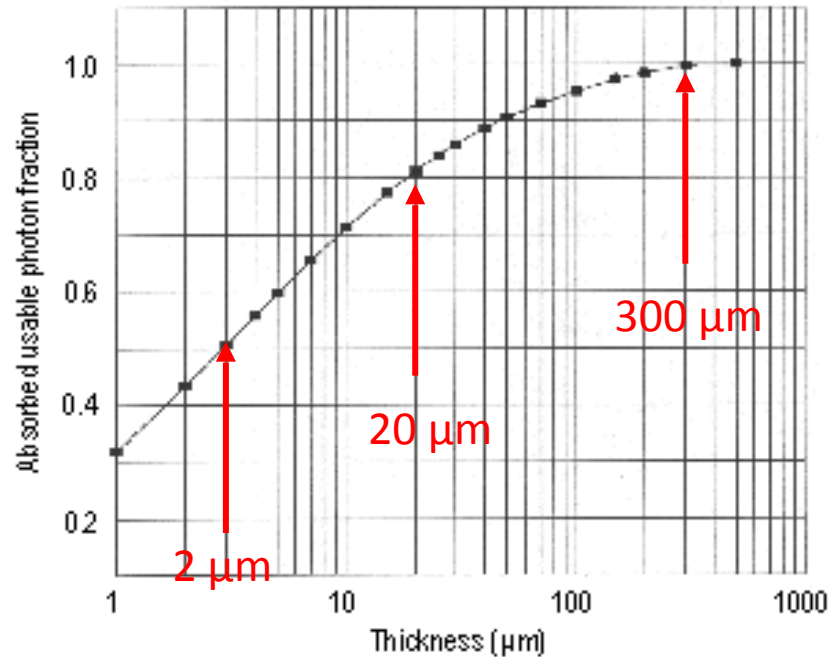
Incident Photon-to-Current Conversion Efficiency (IPCE) or External Quantum Efficiency (EQE) =

$$\frac{\text{Collected electrons at a given wavelength}}{\text{Photons in at a given wavelength}} = \frac{J_{sc}/q}{P_{in}/h\nu} = \frac{J_{sc}(A/cm^2)}{P_{in}(W/cm^2)} \times \frac{1240}{\lambda(nm)} \times 100$$

Absorbed Photon-to-Current Conversion Efficiency (APCE) or Internal Quantum Efficiency (IQE):

$$\frac{\text{Collected electrons at a given wavelength}}{\text{Absorbed photons at a given wavelength}} = \frac{EQE}{1 - R - T} = \frac{EQE}{LHE} \approx \frac{EQE}{1 - 10^{-A}}$$

LHE = Light harvesting efficiency



Ex) While 100% of the light is absorbed by a 300 μm Si wafers, this falls off to 80% in a 20 μm thin layers and 50% in a 2 μm thin layers.

1. IQE = 100% and 2 μm Si → EQE = 50%
2. IQE = 50% and 300 μm Si → EQE = 50%
3. IQE = 100% and 300 μm Si → EQE = 100%

IQE is determined by charge management, e.g. charge separation/collection.

EQE involves both charge management and light management.

Incident Photon-to-Current Conversion Efficiency (IPCE) or External Quantum Efficiency (EQE) =

$$\frac{\text{Collected electrons at a given wavelength}}{\text{Photons in at a given wavelength}} = \frac{J_{sc}/q}{P_{in}/h\nu} = \frac{J_{sc}(A/cm^2)}{P_{in}(W/cm^2)} \times \frac{1240}{\lambda(nm)} \times 100$$

Absorbed Photon-to-Current Conversion Efficiency (APCE) or Internal Quantum Efficiency (IQE):

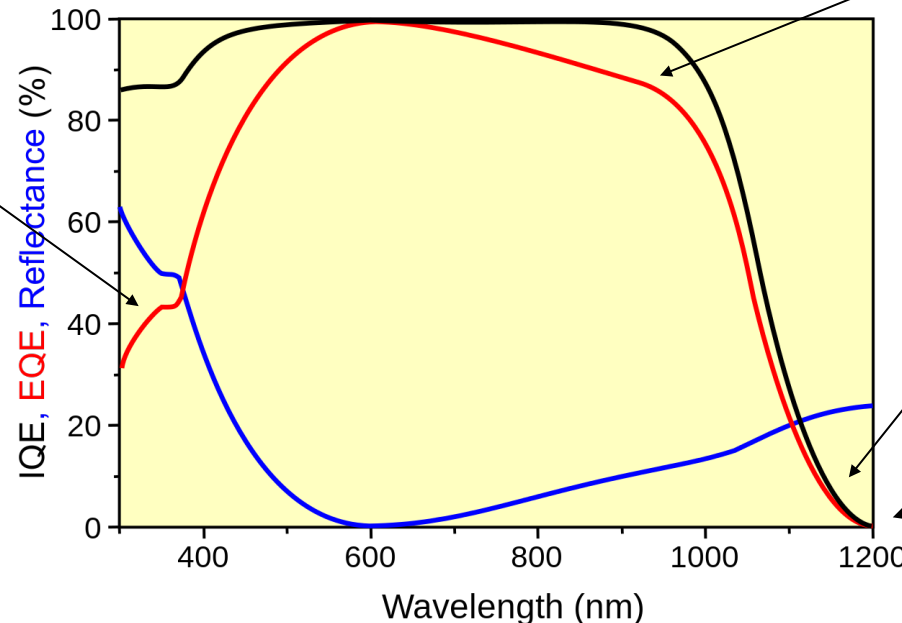
$$\frac{\text{Collected electrons at a given wavelength}}{\text{Absorbed photons at a given wavelength}} = \frac{EQE}{1 - R - T} = \frac{EQE}{LHE} \approx \frac{EQE}{1 - 10^{-A}}$$

LHE = Light harvesting efficiency

Reflection, Parasitic absorption,
front surface recombination

— Internal Quantum Efficiency
— External Quantum Efficiency
— Surface Reflectance

Reflection, limited diffusion length



Reduced absorption, rear surface
recombination, low diffusion
length

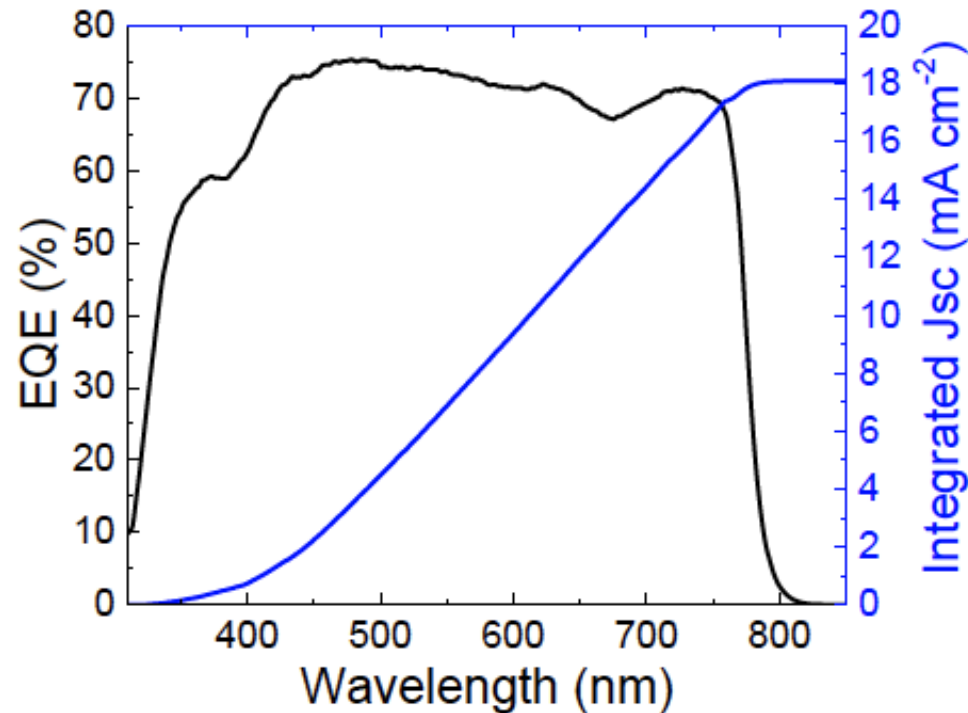
No light is absorbed below the
band gap

Blue: light management
Organe: charge management

When the EQE is integrated overall all wavelength, it will return J_{sc}

$$\frac{\text{Collected electrons at a given wavelength}}{\text{Photons in at a given wavelength}} = \frac{J_{sc}/q}{P_{in}/h\nu}$$

$$J_{sc} = q \int_{\lambda_1}^{\lambda_2} EQE(\lambda) \Phi^{AM1.5} d\lambda \quad \Phi^{AM1.5} : \text{Spectral photon flux}$$



From C. Perini et al., *J. Mater. Chem. A*, 8, 25283 – 25289 (2020)

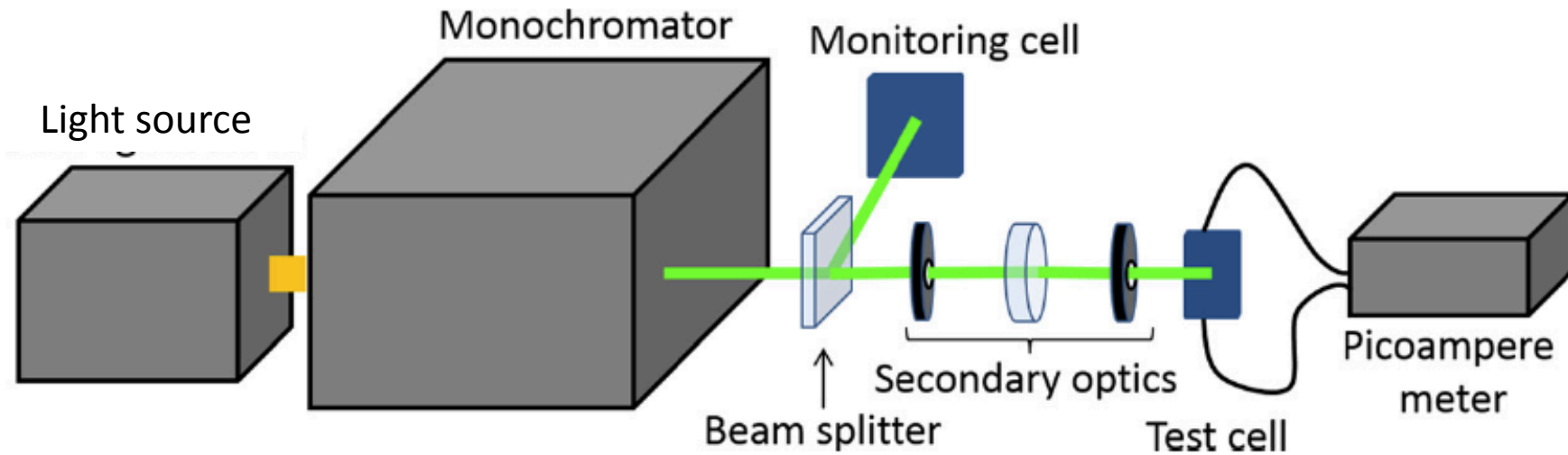


Image from Leyre *et al.*, *J. Renewable Sustainable Energy* 7, 043130 (2015)

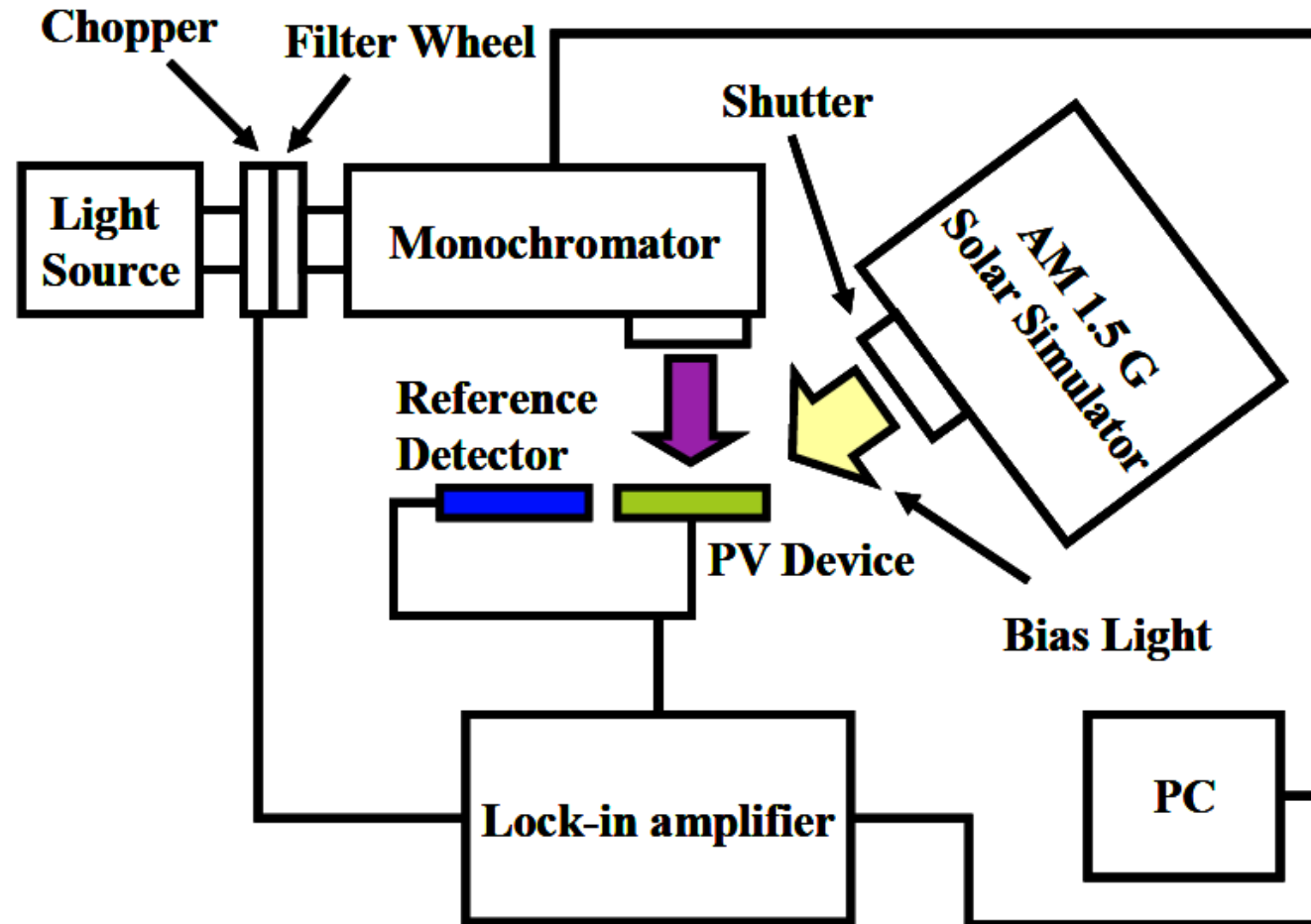


Image from C-F Lin *et al.*, *Int. J. Mol. Sci.*, 12, 476-505 (2011)